Package 'spectral'

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Type Package

Title Common Methods of Spectral Data Analysis

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Description On discrete data spectral analysis is performed by Fourier and Hilbert transforms as well as with model based analysis called Lomb-Scargle method. Fragmented and irregularly spaced data can be processed in almost all methods. Both, FFT as well as LOMB methods take multivariate data and return standardized PSD. For didactic reasons an analytical approach for deconvolution of noise spectra and sampling function is provided.

A user friendly interface helps to interpret the results.

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```
.onAttach
```

Setting up multithread BLAS library

Description

The number of used cores is set to RhpcBLASctl::get_num_cores() on the attach event of the package.

Usage

```
.onAttach(libname, pkgname)
```

Arguments

libname	a character string giving the library directory where the package defining the namespace was found.
pkgname	a character string giving the name of the package.

Details

The algebra system of R relies on a BLAS library which can be set to use many threads / cores. This feature is considered as experimental since there are many differences across the operating systems R is running on. If there is an issue and there is a need to run R in multi-thread mode, consider to install a different optimized version of BLAS. If necessary, the number of cores required can also be changed manually by calling blas_set_num_threads(nCores) and omp_set_num_threads(nCores).

This function is invoked automatically

.onDetach

Description

This function is invoked automatically

Usage

```
.onDetach(libpath)
```

Arguments

libpath a character string giving the complete path to the package.

amax

Local Maxima

Description

Determines all local maxima from a real valued vector.

Usage

amax(x)

Arguments ×

numeric vector

Details

The purpose is to detect all local maxima in a real valued 1D vector. If the first element x[1] is the global maximum, it is ignored, because there is no information about the previous element. If there is a plateau, the first edge is detected.

Value

returns the indicies of local maxima. If x[1] = max, then it is ignored.

Examples

a <- c(1,2,3,2,1,5,5,4) amax(a) # 3, 6 analyticFunction Analytic function

Description

In general a causal real valued signal in time has negative frequencies, when a Fourier transform is applied. To overcome this, a complex complement can be calculated to compensate the negative frequency spectrum. The result is called analytic signal or analytic function, which provides a one sided spectrum.

Usage

```
analyticFunction(x)
```

Arguments

х

real valued data vector

Details

An analytic function xa is composed of the real valued signal representation y and its Hilber transform H(y) as the complex complement

$$xa(t) = x(t) + iH(x(t))$$

. In consequence, the analytic function has a one sided spectrum, which is more natural. Calculating the discrete Fourier transform of such a signal will give a complex vector, which is only non zero until the half of the length. Every component higher than the half of the sampling frequency is zero. Still, the analytic signal and its spectrum are a unique representation of the original signal x(t). The new properties enables us to do certain filtering and calculations more efficient in the spectral space compared to the standard FFT approach. Some examples are:

- **Filtering** because the spectrum is one sided, the user must only modifive values in the lower half of the vector. This strongly reduces mistakes in indexing. See filter.fft
- **Envelope functions** Since the Hilbert transform is a perfect phase shifter by pi/2, the envelope of a band limited signal can be calculated. See envelope
- **Calculations** Deriving and integrating on band limited discrete data becomes possible, without taking the symmetry of the discrete Fourier transform into account. The second example of the spec.fft function calculates the derivative as well, but plays with a centered spectrum and its corresponding "true" negative frequencies

A slightly different approach on the analytic signal can be found in R. Hoffmann "Signalanalyse und -erkennung" (Chap. 6.1.2). Here the signal x(t) is split into the even and odd part. According to Marko (1985) and Fritzsche (1995) this two parts can be composed to the analytic signal, which lead to the definition with the Hilbert transform above.

Value

Complex valued analytic function

References

R. Hoffmann, Signalanalyse und -erkennung: eine Einfuehrung fuer Informationstechniker, Berlin; Heidelberg: Springer, 1998.

H. Marko, Systemtheorie: Methoden und Anwendungen fuer ein- und mehrdimensionale Systeme.3. Aufl., Berlin: Springer, 1995.

G. Fritzsche, Signale und Funktionaltransformationen - Informationselektronik. Berlin: VEB Verlag Technik, 1985

ΒP

Simple bandpass function

Description

This function represents a simple weightening procedure for spectral filtering accoring to the type ("poly", "sinc", "bi-cubic", "gauss") provided.

Usage

BP(f, fc, BW, n = 3, type = "poly")

Arguments

f	vector of frequencies
fc	center frequency
BW	<pre>bandwidth, with w[abs(f -fc) > BW] == min</pre>
n	degree of the polynom, n can be real, e.g. n = 1.6 as sinc alike
type	Type of weightening function: "poly", "sinc", "bi-cubic", "exp", can be abbrevi- ated

Details

The band pass is represented troughout a function in the form of four different types:

1. polynominial function

$$w = 1 - |((f - fc)/BW)|^n$$

with the degree n. The parameter fc controlls the center frequency and desired band width BW. Outside the band width

$$|f - fc| > BW$$

the result is forced to zero. With n = 1.6 a quasi sinc-filter without side bands can be constructed. A quasi rectangular window can be gained by setting n > 5.

2. sinc function corresponds to a rectangular observation window in time domain with

 $\Delta T \ 1/BW$

. It values ALL frequencies according to the si(x) function. Calculation speed might be reduced.

3. bi-cubic encounters 2nd order interpolation kernel, providing a quasi rectangular observation window.

4. exponential Gauss curve. Here the band width is defined as the value of 90

Value

This function returns a weight vector [0..1], which is to apply to the frequency vector f in a top level function

Examples

```
f <- seq(-50,50,by = 1e-2)
fc <- 0.3
BW <- 0.75
par(mfrow = c(2,1))
curve(BP(x, fc = fc, BW = BW, type = "p"), -2, 2, ylim = c(-0.2, 1)
      ,main = "Filter weights"
      ,xlab = "fx",ylab = "w"
)
curve(BP(x,fc = fc, BW = BW, type = "s"), add = TRUE, lty = 2)
curve(BP(x,fc = fc, BW = BW, type = "b"), add = TRUE, lty = 3)
curve(BP(x,fc = fc, BW = BW, type = "g"), add = TRUE, lty = 4)
abline(v = c(fc,fc+BW,fc-BW), lty = 3, col = "grey")
# the corresponding Fourier-Transforms
ty <- c("p","s","b","g")
A0 <- integrate(BP,fc = fc, BW = BW, type = "s",lower = -2,upper = 2)$value
plot(NA, NA, xlab = "x", ylab = "|A|"
     ,main = "corresponding convolution kernels"
     ,xlim = 2*c(-1,1),ylim = c(0, sqrt(2)*A0/(length(f)*BW*min(diff(f))) )
)
for(i in 1:length(ty))
{
  FT <- spec.fft(y = BP(f,fc,BW,type = ty[i]))</pre>
  lines(FT$fx * length(FT$fx) / diff(range(f)),Mod(FT$A),lty = i)
}
```

deconvolve

Description

The function removes the probable alias peaks in the power spectral density. These projections originate from correlated gaps, missing values and interactions with noise. The function should be considered as *experimental* but with didactic background.

Usage

deconvolve(x, y, SNR.enable = T, SNR.level = 1)

Arguments

x	sampling instances
У	values
SNR.enable	binary value, include or exclude the noise
SNR.level	theshold in the sense of a multiple of mean() noise level

Details

In the special case of a non complete equidistant grid containing the data and missing values (NA), this function performs the deconvolution of Y = fft(y) from the sampling spectrum of the aquisition series x. The data is assumed to exist on a equidistant grid with missing values and gaps.

Given a one dimensional vector y of data this function reverses the spectral convolution of Y = S*X+N, if * describes the convolution operation and Y = F(y) denotes the discrete Fourier transform via the operator F(.). If, the sampling series x is considered to be purely deterministic, which should be the case for captured data, or the distortions (missing values, gaps) are *correlated* (see example), then there exists an analytic inversion of the convolution. Given the general definition of power spectral density $|Y|^2 = |S*X+N|^2$ the challenge is to prove $|S*X+N|^2 |S|^2*|X|^2+|N|^2$. Here N describes a stochastic term of gaussian noise. This issue is solved in correlation space, where convolution becomes a multiplication. The auto correlation function (acf) of y is given by $Ry = F(|Y|^2)$. As a remark, IF we consider the special case of equispaced sampling, modeled by the Diraq distribution $\delta(x)$, it is easy to show that the correlation function of a product is the product of individual correaltation functions, $F(|S*X|^2) = F(|S|^2).F(|X|^2)$.

The aim is, to approximate S as the "true" spectrum. To the cost of the phase information, the result is the standardized power spectral density. The spectral noise term F(N) is approximated by a theshold in Fourier space. Here SNR.level sets the factor of mean(fft(y)) below which noise level is assumed. Above this value, the signal should be present. As a parameter to play with, SNR.enable enables or disables the noise term. This parameter was introduced to be consistent with present approaches, not considering the presence of noise.

Value

list of frequency f and spectral density function S

Examples

```
### Deconvolution ###
#
#
# we define a test function with gaps and noise
# we show:
# - the aliased Fourier spectrum and for comparison Lomb Spectrum
# - the corrected spectrum
#
## definition of the sampling series
x <- seq(0,pi/2,by = 5e-3)</pre>
n <- rnorm(length(x), sd = 0.1)
## definition of the test function
## with 2 frequencies
yf <- function(x)</pre>
{
  cos(2*pi*5.123*x) +
  cos(2*pi*17*x)
}
y \leq yf(x)
y <- y - mean(y)
## define strongly correlated gaps
i <- NULL
i <- c(i,which(sin(2*pi / 0.3 * x) - 0.5 > 0))
i <- c(i,which(sin(2*pi / 0.04 * x + 1.123) - 0.5 > 0))
i <- sort(unique(i))</pre>
xs <- x
ys <- yf(xs) + n # add some noise
ys[i] <- NA
## for comparison we calculate a Lomb-Spectrum
LT <- spec.lomb(x = xs, y = ys)
                 f = seq(0, 250, by = 0.02)
                 ,mode = "generalized"
)
WS <- deconvolve(x = xs, y = ys, SNR.enable = 1, SNR.level = 1)
FT <- spec.fft(x = xs, y = ys,center = FALSE)</pre>
FTS <- spec.fft(x = xs, y = is.na(ys),center = FALSE)</pre>
LTS <- spec.lomb(x = xs, y = is.na(ys), f = seq(0, 250, by = 0.02))
### results ###
#
# - signal spectrum (solid) dominant peaks at around f0 = \{5, 17\}
# - (minor) alias peaks (grey line, FFT dots) at f0 +/- fs
```

envelope

```
# - sampling spectrum (dashed) with fs = {3.3, 25} (dominant modes)
# - deconvolved spectrum (solid black) rejects the aliases and sampling
#
#
### time series
par(mfrow = c(1,1), mar = c(4,4,3,0.3))
curve(yf,0,max(x), col = "grey",n = 1000
      xlim = c(0, max(x)), ylim = c(-2, 3)
      ,xlab = "Time", ylab = "y(t)"
      ,main = "Fragmented Time Series"
      )
points(xs,ys)
points(xs[is.na(ys)],yf(xs[is.na(ys)]),pch = 16,cex = 0.5)
legend("topright",c("y(t)","y(tn) + n(tn)","NA's")
       , lty = c(1, NA, NA)
       ,lwd = c(1,NA,NA)
       ,pch = c(NA, 1, 16)
       ,col = c("darkgrey","black","black")
       ,bg = "white"
       , cex = 0.8
)
## plot spectra
par(mfrow = c(1,1), mar = c(4,4,3,0.3))
with(FT,plot(fx,PSD,type="p",log = "x"
     # ,col="grey"
     ,xlim = c(1,100),ylim = c(1e-2,0.75)
     ,xlab = "f", ylab = "PSD"
     ,pch = 1
     ,1wd = 1
     ,main = "Spectra"
))
with(LT,lines(f,PSD,col = "grey",lwd = 4))
with(WS,lines(f,S, lwd = 2, col = "black"))
with(LTS,lines(f,PSD,lty = 2))
abline(h = c(1, 0.5), lty = 3)
legend("topright",c("Fourier","Lomb","Decon.","Sampling")
       ,lty = c(NA,1,1,2)
       ,1wd = c(2,2,2,2)
       ,pch = c(1,NA,NA,NA)
       ,col = c("black","grey","black","black")
       ,bg = "white"
       , cex = 0.8
       ,ncol = 2
```

```
)
```

envelope

Calculates the envelope of a band limited signal

Description

The envelope of an amplitude modulated signal can be calculated by using the Hilbert transform H(y) of the signal or the analytic signal.

Usage

envelope(y)

Arguments

у

numeric vector of the signal

Details

An amplitude modulated function $y(x) = A(x) * cos(\omega * x)$ can be demodulated as follows:

$$A(x)^{2} = y(x)^{2} + H(y(x))^{2}$$

If the signal is not band limited, strange things can happen. See the ripple at the edges in the example below. Pay attention, that the envelope is always the real part of the returned value.

Value

real valued envelope function of the signal

Examples

```
## noisy signal with amplitude modulation
x <- seq(0,1, length.out=2e2)
# original data
y <- (abs(x-0.5))*sin(20*2*pi*x)
ye <- base::Re(envelope(y))
# plot results
plot(x,y,type="1",lwd=1,col="darkgrey",lty=2,ylab="y",main="Spectral filtering")
lines(x,ye)
legend("bottomright",c("modulated","envelope"),col=c("grey","black"),lty=c(2,1))
```

filter.fft

Description

This function provides a method to band pass filter in the frequency domain.

Usage

```
filter.fft(
  y = stop("y-value is missing"),
  x = NULL,
  fc = 0,
  BW = 0,
  n = 3,
  type = "poly"
)
```

Arguments

У	numeric data vector
х	optional x-coordinate
fc	center frequency of the bandpass
BW	bandwith of the bandpass
n	parameter to control the stiffness of the bandpass
type	type of weightening function: "poly", "sinc", "bi-cubic", "gauss", can be abbreviated

Details

A signal y is meant to be equally spaced and causal, which means it starts at t = 0. For times y < 0 the signal is not defined. The filtering itself takes place with the analytic function of y which provides an one sided spectrum. Applying the Fourier transform, all properties of y will be preserved.

The band pass is represented throughout a function in the form of four different types, i.e. "polynom", " $\sin(x)/x$ ", "bi-cubic", "gauss". A detailed description about these types can be found in BP.

Setting fc = 0 one can achieve a low pass filter.

Examples

```
## noisy signal with amplitude modulation
x <- seq(0,1, length.out=500)</pre>
```

original data

filter.lomb

Filter and reconstruction of data analysed via spec.lomb

Description

Given an object of class lomb, this function allows the reconstruction of the input signal using (a) a frequency selection of single or multiple frequency (ranges), and/or (b) the most significant peaks in the periodogram.

Usage

```
filter.lomb(
    l = stop("No Lomb-Data"),
    newx = NULL,
    threshold = 6,
    filt = NULL,
    phase = "nextnb"
)
```

Arguments

1	lomb object
newx	vector of new values at which the restored function is to be evaluated
threshold	statistical threshold in terms of a standard deaviation of the amplidudes. It de- termines which frequencies are used. Lower values give more frequencies.
filt	vector or matix of frequencies (ranges) in which to select the frequencies
phase	set the method to determine the phase at a given frequency

gLmb

Details

To properly reconstruct the signal out of the calculated lomb-object, three different methods are available, which are controlled by the filt-argument.

- 1. If filt=NULL, the most significant values in the (dense) spectrum are used.
- If filt=c(f1,..,fn), the given frequencies are used. The corresponding phase is approximated.
- 3. If class(filt)=="matrix", each row of the 2 x n matrix defines a frequency range. With in each range the "significant" frequencies are selected for reconstruction.

Prior to the reconstruction the filter.lomb-function calculates the most significant amplitudes and corresponding phases. As a measure to select the "correct" frequencies, the threshold argument can be adjusted. The corresponding phases of the underlying sine/cosine-waves are estimated by one of the four following methods.

- 1. phase=="nextnb"... use the phase of the bin of nearest neighbour.
- 2. phase=="lin"... linear interpolation between the two closest bins.
- 3. phase=="lockin"... principle of lock-in amplification, also known as quadrature-demodulation technique.
- 4. phase=="fit"... non-linear least squares fit with stats::nls

Value

This function returns a list which contains the reconstruction according to the lomb-object and news for the given data x and y. The returned object contains the following:

x, y reconstructed signal

f, A, phi used parameters from the lomb-object

p corresponding significance values

gLmb

generalized Lomb-Scargle estimation function

Description

calculates the generalized Lomb-Scargle estimation after Zechmeister et al. (2009)

Usage

gLmb(f, dat, w, Y, hYY)

Arguments

f	frequency
dat	spatial vector including locations and values
W	vector of weights
Υ	weighted sum of values
hYY	weighted sum of squared values

Details

This method is based on the generalized approach

 $y(t) = a * \cos(w * t) + b * \sin(w * t) + c$

which contains the floating average value c of the model function above. The calculation is vectorized to enhance calculation speed.

Н

The Hilbert transformation

Description

The Hilbert transform is a phase shifter, which represents the complex complement to a real vauled signal. It is calculated in the complex frequency space of the signal by using the Fourier transform. Finally, calculating f = y + i * H(y) gives the analytic signal, with a one sided spectrum. (See analyticFunction)

Usage

H(x)

Arguments

х

real valued time series

Value

A numeric real valued vector is returned

interpolate.fft interpolates data using the Fourier back transform

Description

There are two ways to interpolate data from a given spectrum. Frist, one can do zero padding to cover n new data points. Or, secound the complex amplitude with the associated frequency is taken and evaluated at given points xout. Doing that for all frequencies and amplitudes will give the interpolation. The result is compared to linear approximation for didactic reasons.

Usage

```
interpolate.fft(y, x = NULL, n = NULL, xout = NULL)
```

lmb

Arguments

У	numeric data vector to be interpolated
х	numeric data vector with reference points
n	number of new points
xout	a vector new points

Value

A list with a x and y component is returned. The e99 value evaluates the error of the interpolation with respect to linear approximation with the approx() function.

lmb

Lomb-Scargle estimation function

Description

calculates the standard Lomb-Scargle estimation. The calculation is vectorized to enhance calculation speed.

Usage

lmb(f, dat, var_val)

Arguments

f	frequency
dat	spatial vector including locations and values
var_val	variance of the data

plot.fft

Plot fft-objects

Description

This is a wrapper function to plot fft-class objects.

Usage

S3 method for class 'fft'
plot(x, ...)

Arguments

х	Object of the class fft
	further arguments to the plot functions

See Also

spec.fft

Examples

See spec.fft

plot.lomb

plot method for Lomb-Scargle periodograms

Description

This method plots a standard Lomb-Scargle periodogram, which contains the normalized power spectra PSD and the corresponding false alarm probability p. For more details refer to Zechmeister et al. (2009).

Usage

```
## S3 method for class 'lomb'
plot(
  х,
  FAPcol = 1,
  FAPlwd = 1,
  FAPlty = "dashed",
  FAPlim = c(1, 0.001),
  FAPlab = "FAP",
  legend.pos = "topleft",
  legend.cex = 1,
  legend.on = T,
  legend.text = c("Spectrum", "False Alarm Probability"),
  legend.lwd = NULL,
  legend.lty = NULL,
  legend.col = NULL,
  xlab = "Frequency",
 ylab = "Normalized PSD",
 main = "",
  . . .
)
```

Arguments

х	object of class lomb
FAPcol	color of the FAP line
FAPlwd	line width of the FAP line
FAPlty	line type for the FAP graph
FAPlim	limits to the FAP

print.fft

FAPlab	label of the right vertical axis
legend.pos	position of the legend
legend.cex	cex value for the legend
legend.on	logical, wheater to draw a legend or not
legend.text	legend text
legend.lwd	line width
legend.lty	line type
legend.col	color vector of the legend elements
xlab	a label for the x axis, defaults to a description of x.
ylab	a label for the y axis, defaults to a description of y.
main	setting the title of the plot
	further parameters to the plot function

Details

The plot.lomb function is a wrapper function for R's standard scatter plot To switch off certain properties, simply overwrite the parameter. For example log = "" will reset the plot axis back to non-log scale.

References

M. Zechmeister and M. Kurster, "The generalised Lomb-Scargle periodogram. A new formalism for the floating-mean and Keplerian periodograms", Astronomy & Astrophysics, 496(2), pp. 577–584, 2009.

See Also

spec.lomb

Examples

See spec.lomb

print.fft

FFT-Plotting Function

Description

It calls the summary function.

Usage

S3 method for class 'fft'
print(x, ...)

Arguments

х	lomb object
	not used

Value

This function returns nothing

Examples

see summary.lomb() function

print.lomb

Lomb-Plotting Function

Description

It calls the summary function.

Usage

S3 method for class 'lomb'
print(x, ...)

Arguments

х	lomb object
	not used

Value

This function returns nothing

Examples

see summary.lomb() function

Description

This function calculates the Fourier spectrum and power spectral density of a given data object. The dimension of the array can be of arbitary size e. g. 3D or 4D.

Usage

spec.fft(y = NULL, x = NULL, z = NULL, center = T)

Arguments

У	1D data vector, y coordinate of a 2D matrix, nD (even 2D) array or object of class fft
x	x-coordinate of the data in y or z. If y is an array, x must be a named list $x = list(x =, y =)$.
z	optional 2D matrix
center	logical vector, indicating which axis to center in frequency space

Details

The function returns an user friendly object, which contains as much frequency vectors as ordinates of the array. spec.fft provides the ability to center the spectrum along multiple axis. The amplitude output is already normalized to the sample count and the frequencies are given in terms of $1/\Delta x$ -units.

Value

An object of the type fft is returned. It contains the spectrum A, with "reasonable" frequency vectors along each ordinate. psd represents the standardized power spectral density, [0,1]. The false alarm probability (FAP) p is given similar to the Lomb-Scargle method, see spec.lomb.

Missing Values

Given a regular grid $x_i = \delta x \cdot i$ there might be missing values marked with NA, which are treated by the function as 0's. This "zero-padding" leads to a loss of signal energy being roughly proportional to the number of missing values. The correction factor is then (1 - Nna/N) as long as Nna/N < 0.2. If the locations of missing values are randomly distributed the implemented procedure workes quite robust. If correalted gaps are present, the proposed correction is faulty and scales wrong. This is because a convolution of the incomplete sampling vector with the the signal takes place. An aliasing effect takes place distorting the spectral content.

To be compatible with the underlying Fourier transform, the amplitudes are not affected by this rescaling. Only the power spectral density (PSD) is corrected in terms of the energy content, which is experimental for the moment.

See Also

plot.fft

Examples

```
# 1D Example with two frequencies
x <- seq(0, 1, length.out = 1e3)</pre>
y <- sin(4 * 2 * pi * x) + 0.5 * sin(20 * 2 * pi * x)
FT <- spec.fft(y, x)</pre>
par(mfrow = c(2, 1))
plot(x, y, type = "1", main = "Signal")
plot(
 FT,
 ylab = "Amplitude",
 xlab = "Frequency",
 type = "1",
 xlim = c(-30, 30),
 main = "Spectrum"
)
summary(FT)
# 2D example with a propagating wave
x <- seq(0, 1, length.out = 50)</pre>
y <- seq(0, 1, length.out = 50)</pre>
# calculate the data
m <- matrix(0, length(x), length(y))</pre>
for (i in 1:length(x))
 for (j in 1:length(y))
   m[i, j] <- sin(4 * 2 * pi * x[i] + 10 * 2 * pi * y[j])</pre>
# calculate the spectrum
FT <- spec.fft(x = x, y = y, z = m)
# plot
par(mfrow = c(2, 1))
rasterImage2(x = x,
          y = y,
          z = m,
          main = "Propagating Wave")
plot(
 FT,
 main = "2D Spectrum",
 palette = "wb"
 xlim = c(-20, 20),
 ylim = c(-20, 20),
 zlim = c(0, 0.51)
```

spec.fft

```
xlab = "fx",
  ylab = "fy",
  zlab = "A",
  ndz = 3,
  z.adj = c(0, 0.5)
 z.cex = 1
)
summary(FT)
# 3D example with a propagating wave
# sampling vector
x <- list(x = seq(0,2,by = 0.1)[-1]</pre>
          y = seq(0,1, by = 0.1)[-1]
          z = seq(0,1, by = 0.1)[-1]
)
# initializing array
m <- array(data = 0,dim = sapply(x, length))</pre>
for(i in 1:length(x$x))
  for(j in 1:length(x$y))
    for(k in 1:length(x$z))
      m[i,j,k] <- cos(2*pi*(1*x$x[i] + 2*x$y[j] + 2*x$z[k])) + sin(2*pi*(1.5*x$x[i]))^2</pre>
FT <- spec.fft(x = x, y = m, center = c(TRUE,TRUE,FALSE))</pre>
par(mfrow = c(2,2))
# plotting m = 0
rasterImage2( x = FT$fx
             ,y = FT$fy
              ,z = abs(FT$A[,,1])
              ,zlim = c(0,0.5)
              ,main="m = 0"
              )
# plotting m = 1
rasterImage2( x = FT$fx
             ,y = FT$fy
              ,z = abs(FT$A[,,2])
              ,zlim = c(0,0.5)
              ,main="m = 1"
)
\# plotting m = 2
rasterImage2( x = FT$fx
             ,y = FT$fy
              ,z = abs(FT$A[,,3])
              ,zlim = c(0,0.5)
              ,main="m = 2"
```

```
)
rasterImage2( x = FT$fx
             ,y = FT$fy
              ,z = abs(FT$A[,,4])
              ,zlim = c(0,0.5)
              ,main="m = 3"
)
summary(FT)
# calculating the derivative with the help of FFT
*****
#
# Remember, a signal has to be band limited.
# !!! You must use a window function !!!
#
# preparing the data
x <- seq(-2, 2, length.out = 1e4)</pre>
dx <- mean(diff(x))</pre>
y <- win.tukey(x) * (-x ^ 3 + 3 * x)
# calcualting spectrum
FT <- spec.fft(y = y, center = TRUE)</pre>
# calculating the first derivative
FT$A <- FT$A * 2 * pi * 1i * FT$fx
# back transform
dm <- spec.fft(FT)</pre>
# plot
par(mfrow=c(1,1))
plot(
  х,
  c(0, diff(y) / dx),
  type = "1",
  col = "grey",
  lty = 2,
  ylim = c(-4, 3)
)
# add some points to the line for the numerical result
points(approx(x, Re(dm$y) / dx, n = 100))
# analytical result
curve(-3 * x ^ 2 + 3,
      add = TRUE,
      1ty = 3,
      n = length(x)
legend(
  "topright",
  c("analytic", "numeric", "spectral"),
  title = "diff",
  lty = c(3, 2, NA),
```

spec.lomb

```
pch = c(NA, NA, 1),
col=c("black","grey","black")
)
title(expression(d / dx ~ (-x ^ 3 + 3 * x)))
```

spec.lomb

Lomb-Scargle Periodigram

Description

The Lomb-Scargle periodigram represents a statistical estimator for the amplitude and phase at a given frequency. This function takes also multivariate (n-dimensional) input data.

Usage

```
spec.lomb(
  x = NULL,
  y = stop("Missing y-Value"),
  f = NULL,
  ofac = 1,
  w = NULL,
  mode = "normal",
  maxMem = 8,
  cl = NULL
)
```

Arguments

x	sampling vector or data frame data.frame(x1,x2,x3,)
У	input data vector or data frame data.frame(x1,x2,,val)
f	optional frequency vector / data frame. If not supplied f is calculated.
ofac	in case f=NULL this value controlls the amount of frequency oversampling.
w	weights for data. It must be a 1D vector.
mode	"normal" calculates the normal Lomb-Scargle periodogram; "generalized" calculates the generalized Lomb-Scargle periodogram including floating average and weights.
maxMem	sets the amount of memory (in MB) to utilize, as a rough approximate.
cl	

Details

Since the given time series does not need to be evenly sampled, the data mainly consists of data pairs x1, x2, x3, ... (sampling points) and (one) corresponding value y, which stores the realisation/measurement data. As can be seen from the data definition above, multivariate (n-dimensional) input data is allowed and properly processed.

Two different methods are implemented: the standard Lomb-Scargle method with

 $y(t) = a * \cos(\omega(t-\tau)) + b * \sin(\omega(t-\tau))$

as model function and the generalized Lomb-Scargle (after Zechmeister 2009) method with

 $y(t) = a * \cos(\omega t) + b * \sin(\omega t) + c$

as model function, which investigates a floating average parameter c as well.

Both methods can be supplied by an artifical dense frequency vector f. In conjunction with the resulting phase information the user might be able to build a "Fourier"-like spectrum to reconstruct or interpolate the timeseries in equally spaced sampling. Remind the band limitation which must be fulfilled for this.

- **f** The frequencies should be stored in a 1D vector or in case of multivariate analysis in a data.frame structure to preserve variable names
- ofac If the user does not provide a corresponding frequency vector, the ofac parameter causes the function to estimate

$$nf = ofac * length(x)/2$$

equidistant frequencies.

p-value The p-value (aka false alarm probability FAP) gives the probability, wheter the estimated amplitude is NOT significant. However, if p tends to zero the amplidutde is significant. The user must decide which maximum value is acceptable, until an amplitude is not valid.

If missing values NA or NaN appear in any column, the corresponding row is excluded from calculation.

Value

The spec.lomb function returns an object of the class lomb, which is a list containg the following information:

- A A vector with amplitude spectrum
- f corresponding frequency vector
- phi phase vector

PSD power spectral density normalized to the sample variance

floatAvg floating average value only in case of mode == "generalized"

- w if, mode == "generalized" contains the weighting vector
- x,y original data
- **p** p-value False Alarm Probability

spec.lomb

Speed Up

In general the function calculates everything in a vectorized manner, which speeds up the procedure. If the memory requirement is more than maxMem, the calculation is split into chunks which fit in the memory (cache). Depending on the problem size (number of frequencies and data size) a tuning of this value enhances speed.

Please consider to replace the BLAS library by a multithreaded version. For example https:// prs.ism.ac.jp/~nakama/SurviveGotoBLAS2/binary/windows/x64/ is hosting some Windows RBlas.dll files. Refer to https://mattstats.wordpress.com/2016/02/07/r-with-gotoblas-on-windows-10/ for further information.

The parameter cl controls a possible cluster, which can be invoked. It takes an integer number of workers (i. e. cl = 4), a list with node names c("localhost", ...) or an object of class 'cluster' or similar. The first two options cause the function to create the cluster internally. This takes time due to the initialization. The faster way is to provide an already initialized cluster to the function.

References

A. Mathias, F. Grond, R. Guardans, D. Seese, M. Canela, H. H. Diebner, and G. Baiocchi, "Algorithms for spectral analysis of irregularly sampled time series", Journal of Statistical Software, 11(2), pp. 1–30, 2004.

J. D. Scargle, "Studies in astronomical time series analysis. II - Statistical aspects of spectral analysis of unevenly spaced data", The Astrophysical Journal, 263, pp. 835–853, 1982.

M. Zechmeister and M. Kurster, "The generalised Lomb-Scargle periodogram. A new formalism for the floating-mean and Keplerian periodograms", Astronomy & Astrophysics, 496(2), pp. 577–584, 2009.

See Also

filter.lomb

Examples

```
# create two sin-functions
x_orig <- seq(0,1,by=1e-2)
y_orig <- 2*sin(10*2*pi*x_orig) + 1.5*sin(2*2*pi*x_orig)
# make a 10% gap
i <- round(length(x_orig)*0.2) : round(length(x_orig)*0.3)
x <- x_orig
y <- y_orig
x[i] <- NA
y[i] <- NA
# calculating the lomb periodogram
l <- spec.lomb(x = x, y = y,ofac = 20,mode = "normal")
# select a frequency range
m <- rbind(c(9,11))
# select and reconstruct the most significant component
```

```
12 = filter.lomb(l, x_orig, filt = m)
# plot everything
par(mfrow=c(2,1),mar = c(4,4,2,4))
plot(x,y,"l", main = "Gapped signal")
lines(12$x, 12$y, 1ty=2)
legend("bottomleft",c("gapped","10Hz component"),lty=c(1,2))
plot(1,main = "Spectrum")
summary(1)
### Multivariate -- 3D Expample ###
require(lattice)
fx <- 8.1
fy <- 5
fz <- 2
# creating frequency space
f <- expand.grid( fx = seq(-10,10,by = 0.5)</pre>
                  , fy = seq(-10, 10, by = 0.5)
                  , fz = 0:3
)
# creating spatial space
pts <- expand.grid( x = seq(0, 1, by = 0.02)
                   ,y = seq(0,1,by = 0.02)
                    z = seq(0, 1, by = 0.02)
)
# gapping 30%
i <- sample(1:dim(pts)[1],0.7*dim(pts)[1])</pre>
pts <- pts[i,]</pre>
# caluculating function
pts$val <- cos(2*pi*( fx*pts$x</pre>
                     + fy*pts$y
                     + fz*pts$z
                     ) + pi/4
              ) +
  0.5 * cos(2*pi*( - 0.5 * fx*pts$x
              + 0.5*fy*pts$y
              + 1 * pts$z
  ) + pi/4
  )
# display with lattice
levelplot(val~x+y,pts,subset = z == 0,main = "with z = 0")
# calculating lomb takes a while
# or we sample only a few points
# which enlarges the noise but accelerates the calculation
1 <- spec.lomb(y = pts[sample(1:dim(pts)[1],2e3),]</pre>
```

summary.fft

```
,f = f
# ,mode = "generalized"
)
# name the stripes
l$fz_lev <- factor(x = paste("fz =",l$fz)
)
# display output
levelplot(PSD~fx+fy|fz_lev,l)
# the result is an oversampled spectrum of a non equidistant
# sampled function. We recognize a 3D analysis in all provided
# spatial directions x, y, z.
summary(l)
```

```
summary.fft
```

Summarize FFT objects

Description

The function summarizes properties from the class(fft) object.

Usage

S3 method for class 'fft'
summary(object, p0 = 0.01, ...)

Arguments

object	lomb object
p0	False Alarm Probability (FAP) threshold, default 1%
	not used

Details

The false alarm probability threshold p0 value can be changed to modify the amount of significant peaks.

Value

a list of significant values of the spectral analysis

Examples

see spec.fft() example

summary.lomb

Description

The function summarizes properties from the Lomb object.

Usage

S3 method for class 'lomb'
summary(object, p0 = 0.01, ...)

Arguments

object	lomb object
p0	False Alarm Probability threshold, default 1%
	not used

Details

The false alarm probability threshold p0 value will adjust the number of peaks.

The effectiveBandWidth describes the coverage of processed frequencies by the spec.lomb function. If the ratio to averageSampling is almost 2, then the Nyquist criterion can be assumed to be fullfilled. If the ratio is much less than 2 then only a fraction of information is analysed.

The minFreqStep is an estimate of the minimum frequency step determined from the Lomb-Object.

Average sampling is calculated from the median distance between two spatial points.

The possible frequency resolution originates also from the spatial (temporal) input data by 1/(diff(range(x))), if x is the spatial (temporal) coordinate.

Value

a list of significant values of the spectral analysis

Examples

see spec.lomb() example

waterfall

Description

A waterfall-diagramm displays the local frequency in dependence of or spatial vector. One can then locate an event in time or space.

Usage

```
waterfall(
  y = stop("y value is missing"),
  x = NULL,
  nf = 3,
  type = "b",
  width = 7
)
```

Arguments

У	numeric real valued data vector
x	numeric real valued spatial vector. (time or space)
nf	steepness of the bandpass filter, degree of the polynomial.
type	type of weightening function: "poly", "sinc", "bi-cubic", "gauss", can be abbreviated
width	normalized maximum "inverse" width of the bandpass $bw = fc/width$.

Details

Each frequency is evaluated by calculating the amplitude demodulation, which is equivalent to the envelope function of the band pass filtered signal. The frequency of interest defines automatically the center frequency fc of the applied band pass with the bandwidth BW:

$$BW = fc/width, BW < width - > BW = width, BW > width - > BW = fc/width$$

The frequency is normalized so the minimal frequency is 1. With increasing frequency the bandwidth becomes wider, which lead to a variable resolution in space and frequency. This is comparable to the wavelet (or Gabor) transform, which scales the wavelet (window) according to the frequency. However, the necessary bandwidth is changed by frequency to take the uncertainty principle into account. Slow oscillating events are measured precisely in frequency and fast changing processes can be determined more exact in space. This means for a signal with steady increasing frequency the waterfall function will produce a diagonally stripe. See the examples below.

Value

a special fft-object is returned. It has mode "waterfall" and x and fx present, so it is only plotable.

Missing values

Given a regular grid $x_i = \delta x \cdot i$ there might be missing values marked with NA, which are treated by the function as 0's. This "zero-padding" leads to a loss of signal energy being roughly proportional to the number of missing values. The correction factor is then (1 - Nna/N) as long as Nna/N < 0.2. As long as the locations of missing values are randomly distributed the implemented procedure workes quite robust. If, in any case, the distribution becomes correlated the proposed correction is faulty and projects the wrong energies.

The amplitudes and PSD values are compensated to show up an estimate of the "correct" value. Therefore this method is experimental

Examples

```
#### noisy signal with amplitude modulation ####
x \le seq(0,3, length.out = 1000)
# original data
# extended example from envelope function
y <- 1*(abs(x-1.5))*sin(10*2*pi*x) + ifelse(x > 1.5,sin(15*(1+0.25*(x - 1.5))*2*pi*x),0)
ye <- base::Re(envelope(y))</pre>
par(mfrow=c(2,1),mar=c(1,3.5,3,3),mgp=c(2.5,1,0))
# plot results
plot(x,y,type="l",lwd=1,col="darkgrey",lty=2,ylab="y",main="Original Data",xaxt="n",xlab="")
lines(x, ye)
legend("bottomright",c("modulated","envelope"),col=c("grey","black"),lty=c(2,1))
par(mar=c(3.5,3.5,2,0))
wf <- waterfall(y, x, nf = 3)
# rasterImage2(x = wf$x, y = wf$fx, z = wf$A
               ,ylim = c(0,60))
#
plot(wf,ylim=c(0,40),main="Waterfall")
#### uncertainty principle ####
#
# take a look at the side effects
# at [0,30] and [1,0]
#
# With a large steepness e.g. n = 50 you will gain
# artefacts.
#
# if frequency is not stationary
# PSD becomes > 1 depending on the type of band filter.
#
x <- seq(0,1, length.out=1500)</pre>
y <- sin(100*x*x)</pre>
FT \leq spec.fft(x = x, y = y)
wf <- waterfall(y,x)</pre>
```

```
par(mfrow=c(2,1),mar=c(1,3.5,3,3),mgp=c(2.5,1,0))
# plot results
plot(x,y,type="l",lwd=1,col="darkgrey",lty=2,ylab="y",main="Original Data",xaxt="n",xlab="")
par(mar=c(3.5,3.5,2,0))
plot(wf
             ,ylim=c(0,40),main="Waterfall"
            )
abline(h = 25, lty = 3, lwd = 3, col = "grey")
range(wf$PSD,na.rm = TRUE)
range(wf$A)
###### effect of missing values #####
#
# 10% random missing values cause a
# distortion and a miss scaling of
# the PSD value, which becomes >1 now.
# This depends on the type of band pass
# filter selected.
#
x <- seq(0,5, length.out=500)</pre>
y <- sin(2*pi * 15 * x + 2*1*cos(2*pi*0.5*x))</pre>
# delete 10% of the data
y[sample(length(y), size = 50)] <- NA</pre>
wf <- waterfall(y,x,type = "b")</pre>
par(mfrow=c(2,1),mar=c(1,3.5,3,3),mgp=c(2.5,1,0))
# plot results
plot(x,y,type="l",lwd=1,col="darkgrey",lty=2,ylab="y",main="Original Data",xaxt="n",xlab="")
par(mar=c(3.5,3.5,2,0))
plot(wf
     ,ylim=c(10,20),main="Waterfall"
)
abline(h = 25, lty = 3, lwd = 3, col = "grey")
# check the PSD range
range(wf$PSD)
range(wf$A)
```

```
win.cos
```

Cosine window function

Description

This window function returns a vector of weights with means of a cosine window

win.nutt

Usage

win.cos(n)

Arguments

n

data vector to be windowed

See Also

Windowfunctions

win.hann

Hanning window function

Description

This window function returns a vector of weights with means of a generlized Hann-window.

Usage

win.hann(n, a = 2)

Arguments

n	data vector to be windowed
а	order of the window, default a = 2

See Also

Windowfunctions

win	n 11	++
WIII	. IIU	LL

Nuttall window function

Description

This window function returns a vector of weights with means of a Nuttall-window.

Usage

win.nutt(n, a = c(0.355768, 0.487396, 0.144232, 0.012604, 0))

Arguments

n	data vector to be windowed
а	coefficients default a = c(0.355768,0.487396,0.144232,0.012604,0)

win.tukey

Details

This window function provides a continuous first derivative everywhere, like the Hann window. Adopted from the idea of Hann this window consists of up to 5 trigonometric polynominial terms, i.e.

 $w_n = a_1 - a_2 \cos(2\pi n/M) + a_3 \cos(4\pi n/M) - a_4 \cos(6\pi n/M) + a_5 \cos(8\pi n/M)$

Different sets of coefficients:

Nuttall(Default)	c(0.355768, 0.487396, 0.144232, 0.012604,0)
Blackman-Nuttall	c(0.3635819, 0.4891775, 0.1365995, 0.0106411,0)
Blackman-Harris	c(0.35875, 0.48829, 0.14128, 0.01168,0)
Flat-Top	c(0.211557895, 0.41663158, 0.277263158, 0.083578947, 0.006947368)

See Also

Windowfunctions

win.tukey

Tukey window function

Description

This window function returns a vector of weights with means of a Tukey-window. In contrast to a cosine window this function is more steep at the beginning and the end. And it is 1 in the middle.

Usage

win.tukey(n, a = 0.5)

Arguments

n	data vector to be windowed
а	width of the rising and falling edge as ratio of the total data length

See Also

Windowfunctions

Windowfunctions Windowfunctions

Description

Some typical windowfunctions are defined below:

Details

win.cos() cosine window

win.tukey() Tukey window

win.hann() Hann window

win.nutt() Nutt window

A window function weights a given dataset in a way, that the new data set is coerced to be periodic. This method reduces the leakage effects of the discrete Fourier transform.

Value

All window functions return a wighting vector with the same length as the provided data vector.

Examples

```
y <- 1:100
y_cos <- y * win.cos(y)
y_tuk <- y * win.tukey(y)
y_han <- y * win.hann(y)
# Plot the original data
plot(y,main="Effect of window functions")
legend("topleft",c("original","cos","tukey","han"),pch=c(1,16,17,18))
points(y_cos,pch=16)
points(y_tuk,pch=17)
points(y_han,pch=18)
```

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